

Abstract accepted for oral presentation in Room 308 of Moscone South at 2:55 PM on 7 December 2011 as Paper SM33B-06 in Session SM33B (Magnetospheric Response to Transient Solar-Wind Features II) at the 2011 fall AGU meeting in San Francisco from 5-9 December (abstract control ID 1203677):

Changes in the High-Latitude Topside Ionospheric Vertical Electron-Density Profiles in Response to Solar-Wind Perturbations During Large Magnetic Storms

Robert F. Benson (robert.f.benson@nasa.gov)

NASA/Goddard Space Flight Center, Geospace Physics Laboratory, Code 673,
Heliophysics Science Division, Greenbelt, MD 20771

Joseph Fainberg (fainberg@ieee.org)

NASA/Goddard Space Flight Center Emeritus, Geospace Physics Laboratory, Code 673,
Heliophysics Science Division, Greenbelt, MD 20771

Vladimir Osherovich (vladimir.osherovich@gmail.com)

CUA/Goddard Space Flight Center, Geospace Physics Laboratory, Code 673,
Heliophysics Science Division, Greenbelt, MD 20771

Vladimir Truhlik (vtr@ufa.cas.cz)

Institute of Atmospheric Physics, Academy of Science Czech Republic,
Praha, Czech Republic

Yongli Wang (yongli.wang-1@nasa.gov)

UMBC/GPHI/Space Weather Laboratory, Code 674,
Heliophysics Science Division, Greenbelt, MD 20771

Becca Arbacher (becca.arbacher@gmail.com)

National Space Club Scholars Summer Intern (Montgomery Blair High School, Silver Spring, Maryland)/Goddard Space Flight Center, Geospace Physics Laboratory, Code 673, Heliophysics Science Division, Greenbelt, MD 20771

The latest results from an investigation to establish links between solar-wind and topside-ionospheric parameters will be presented including a case where high-latitude topside electron-density $Ne(h)$ profiles indicated dramatic rapid changes in the scale height during the main phase of a large magnetic storm ($Dst < -200$ nT). These scale-height changes suggest a large heat input to the topside ionosphere at this time. The topside profiles were derived from ISIS-1 digital ionograms obtained from the NASA Space Physics Data Facility (SPDF) Coordinated Data Analysis Web (CDAWeb). Solar-wind data obtained from the NASA OMNIWeb database indicated that the magnetic storm was due to a magnetic cloud. This event is one of several large magnetic storms being investigated during the interval from 1965 to 1984 when both solar-wind and digital topside ionograms, from either Alouette-2, ISIS-1, or ISIS-2, are potentially available.